Data Assimilation Techniques 101
(and Their Use for Ionospheric Science and Applications)

L. Scherliess

Center for Atmospheric & Space Sciences
Utah State University
Logan, Utah 84322

CEDAR Workshop
Santa Fe
June, 2011
The Daily Weather Forecast is a Product of Data Assimilation
On the One Hand, we have large Quantities of Data

- Different kinds of instruments measuring different quantities (apples and oranges)
- Observations are in different places
- Observations have different cadence and availability
- Observations have different error statistics

Difficult to create coherent Picture
On the Other Hand, we have Mature Theoretical/Numerical Models

- Models contain our ‘knowledge’ of the physics

Uncertain Parameters in Physics-Based Model

- O\(^+\) - O Collision Frequency
- Secondary Electron Production
- Downward Heat Flow
- Chemical Reaction Rates
- External Forcing
- Etc.
Objectives

- Optimally combine Data and the Model to create coherent Picture of the Space Environment

- Solution satisfies the physical laws and ‘agrees’ with the data and the model as best as possible (within their error bounds)
Data Assimilation Tasks

- Develop Physical Model
- Develop Assimilation Algorithm
- Data Acquisition Software
- Data Quality Control
- Executive System
- Validation Software
Brief Historical Background

Data Assimilation in the Atmosphere:
- Initial Attempts started in the 1950s (NWP)

Data Assimilation in the Oceans:
- Began with large scales (mean properties) about 30 yrs ago
- Regional efforts (e.g., Gulf stream) [15-20 yrs ago]
- Produce operational upper ocean now- and forecast.
Data Assimilation in Space Sciences

- Assimilative Mapping of Ionospheric Electrodynamics (AMIE, Richmond and Kamide, 1988)
- Initial Testing of Kalman Filter for Ionospheric Electron Density Reconstructions (Howe et al., 1998)
- Data Assimilation Models for the Ionosphere (late 1990): GAIM models, IDA4D
- Data Assimilation Models for the Thermosphere (Minter et al., Fuller-Rowell et al.)
- Data Assimilation for the Radiation Belts
- Initial Attempts for Solar Data Assimilation
What can we learn from Meteorology?

Data Assimilation Techniques have been used in Meteorology for the last 50 years

- Most Accurate Specifications and Forecast Models are Those that
  Assimilate Measurements into a *Physics-Based* Numerical Model

- Better Predictions are Obtained for the Atmosphere
  - When the Data are Assimilated with a Rigorous Mathematical Approach
Data Assimilation Techniques

- **3-d Var**
  \[ J(\delta x) = 1/2\delta x^T P^{-1}\delta x + 1/2[ H(\delta x + x^b) - y^o]^T R^{-1} [H(\delta x + x^b) - y^o] \]

- **4-d Var**
  \[ J(\delta x) = 1/2\delta x_0^T P^{-1}\delta x_0 + 1/2 \sum_{i=0}^{n} [ H_i(M_{i,o}(x_0)) - y_{i}^o]^T R^{-1} [H_i(M_{i,o}(x_0)) - y_{i}^o] \]

- **Kalman Filter**
  \[
  \begin{align*}
  x^f &= Mx + \eta \\
  P^f &= MPM^T + Q \\
  y^o &= Hx + \varepsilon \\
  K &= P^fH^T (HP^fH^T + R)^{-1} \\
  x^a &= x^f + K(y^o - Hx^f) \\
  P^a &= (I-KH)P^f
  \end{align*}
  \]
Physical Model creates a forecast which is adjusted by the Data to create an ‘analysis’, which serves as the start for the next model forecast. In the analysis the Data Errors and Model Errors are used as weights.
Fundamental Concept of 3D-Var

Start with a forecast or an estimate of the state (background)

‘Best-Guess’ Background → Short-Term Forecast → Analysis → Forecast

Data Collection → Quality Control

Utah State University

GA I M
"Bringing The Pieces Together"
Fundamental Concept of 3D-Var

Minimize the difference between the analysis and a *weighted combination* of

- the background and
- the observations.
The Cost Function

To produce the analysis we want to minimize a “Cost Function” $J$ which consists of:

$$J = J_B + J_O + J_C$$

$J_B$: Weighted fit to the background field

$J_O$: Weighed fit to the observations

$J_C$: Constraint which can be used to impose physical properties (e.g., analysis should satisfy Maxwell’s equations, continuity equation, …)
A typical form for the $J_B$ term is:

$$J_B = (x_A - x_B)^T B^{-1} (x_A - x_B)$$

Where:

$x_A$: Analysis Variable (e.g., Electron Density, Temperature, …)

$x_B$: Background Field, obtained from the Model Forecast

$B$: Background Error Covariance Matrix:

- How good is your Background
- What are covariances between different elements

The background error covariances are only poorly known
The Cost Function, cont.

A typical form for the cost function for the observations is:

\[ J_O = [y - H(x_A)]^T R^{-1} [y - H(x_A)] \]

Where:

- \( y \): Represents all Observations
- \( H \): Forward Operator which maps the Grid Point Values to Observations (can be linear or nonlinear)
- \( R \): Observation Error Covariance Matrix:
  - How good is your data?
  - (also includes the representativeness of the data)
The Cost Function, cont.

The **Physical Properties/Model** were used to:

- Obtain the best possible background field
- To constrain the Analysis

- Cost function and the constraints are not explicitly time dependent

- A temporal model is not necessarily required

- Snapshots
Fundamental Concepts of 4D-Var

4D-Var introduces the temporal dimension to data assimilation

Find a close fit to the data that is consistent with the dynamical model over an extended period of time.

⇒ Find the closest trajectory

\[
J(\delta x) = 1/2 \delta x_0^T P^{-1} \delta x_0 + 1/2 \sum_{i=0}^{n} \left[ H_i(M_{i,o}(x_0)) - y_i^o \right]^T R^{-1} \left[ H_i(M_{i,o}(x_0)) - y_i^o \right]
\]
Fundamental Concepts of 4D-Var

4D-Var introduces the temporal dimension to data assimilation

Find a close fit to the data that is consistent with the dynamical model over an extended period of time.

⇒ Find the closest trajectory

\[
J(\delta x) = \frac{1}{2} \delta x_0^T P^{-1} \delta x_0 + \frac{1}{2} \sum_{i=0}^{n} \left[ H_i(M_{i,o}(x_0)) - y_i^o \right]^T R^{-1} \left[ H_i(M_{i,0}(x_0)) - y_i^o \right]
\]

- Model Error Covariance
- Data Error Covariance
Another Data Assimilation Technique
The Kalman Filter

- M - State Transition Matrix
- P - Model Error Covariance
- y - Data Vector
- R - Observation Error Covariance
- X - Model State Vector
- η - Transition Model Error
- Q - Transition Model Error Covariance
- H - Measurement Matrix
- ε - Observation Error
- K - Kalman Gain

Model Error Covariance

\[ x^f = Mx + \eta \]  \hspace{0.5cm} (1)
\[ P^f = MPM^T + Q \]  \hspace{0.5cm} (2)
\[ y^o = Hx + \epsilon \]  \hspace{0.5cm} (3)
\[ K = P^fH^T(HP^fH^T + R)^{-1} \]  \hspace{0.5cm} (4)
\[ x^a = x^f + K(y^o - Hx^f) \]  \hspace{0.5cm} (5)
\[ P^a = (I - KH)P^f \]  \hspace{0.5cm} (6)

Data Error Covariance
Another Data Assimilation Technique  
The Kalman Filter

- M - State Transition Matrix
- P - Model Error Covariance
- y - Data Vector
- R - Observation Error Covariance
- X - Model State Vector
- η - Transition Model Error
- Q - Transition Model Error Covariance
- H - Measurement Matrix
- ε - Observation Error
- K - Kalman Gain

\[ x^f = Mx + \eta \]  
\[ P^f = MPM^T + Q \]  
\[ y^o = Hx + \epsilon \]  
\[ K = P^fH^T(HP^fH^T + R)^{-1} \]  
\[ x^a = x^f + K(y^o - Hx^f) \]  
\[ P^a = (I - KH)P^f \]
Another Data Assimilation Technique
The Kalman Filter

The Dynamical Model entered the Filter:

- Evolution of the State Vector (make a Forecast)
- Evolution of the Error Covariance Matrix

\[ x^f = Mx + n \]  \hspace{1cm} (1)
\[ P^f = MPM^T + Q \]  \hspace{1cm} (2)
\[ y^o = Hx + \epsilon \]  \hspace{1cm} (3)
\[ K = P^fH^T(HP^fH^T + R)^{-1} \]  \hspace{1cm} (4)
\[ x^a = x^f + K(y^o - Hx^f) \]  \hspace{1cm} (5)
\[ P^a = (I - KH)P^f \]  \hspace{1cm} (6)
Another Data Assimilation Technique

The Kalman Filter

The Dynamical Model entered the Filter:

- Evolution of the State Vector (make a Forecast)
- Evolution of the Error Covariance Matrix

\[ \bf{x^f} = \bf{Mx} + \bf{n} \]  
\[ \bf{P^f} = \bf{MPM^T} + \bf{Q} \]  
\[ \bf{y^o} = \bf{Hx} + \epsilon \]  
\[ \bf{K = P^fH^T(HP^fH^T + R)^{-1}} \]  
\[ \bf{x^a} = \bf{x^f} + \bf{K(y^o - Hx^f)} \]  
\[ \bf{P^a} = (\bf{I - KH})P^f \]

\[ \rightarrow \] Error Covariance Matrix becomes time-dependent and evolves with the same physical model as the state!

This is computationally the most expensive step in the Kalman filter
Example: Tracking of a Rocket with a Kalman Filter

A rocket is flying through space launched from an initial location with an initial velocity.

\[ m \frac{d^2 x}{dt^2} = ma \quad \Rightarrow \quad \frac{dx}{dt} = v \quad \Rightarrow \quad x_{i+1} \approx x_i + v_i \cdot dt \]

\[ \frac{dv}{dt} = a \quad \Rightarrow \quad v_{i+1} \approx v_i + a_i \cdot dt \]

\[ a_{i+1} \approx a_i \]

In Kalman filter we have:

\[ x_{i+1} = M x_i \]

\[
\begin{pmatrix}
  x_{i+1} \\
  v_{i+1} \\
  a_{i+1}
\end{pmatrix} =
\begin{pmatrix}
  1 & dt & 0 \\
  0 & 1 & dt \\
  0 & 0 & 1
\end{pmatrix}
\begin{pmatrix}
  x_i \\
  v_i \\
  a_i
\end{pmatrix}
\]
Example: Tracking of a Rocket with a Kalman Filter

Propagate Error Covariance Matrix: \( P_{i+1} = M P_i M^T \)

\[
P_0 = \begin{pmatrix}
\sigma_x^2 & 0 & 0 \\
0 & \sigma_v^2 & 0 \\
0 & 0 & \sigma_a^2
\end{pmatrix}
\]

At the next time step:

\[
P_1 = \begin{pmatrix}
1 & dt & 0 \\
0 & 1 & dt \\
0 & 0 & 1
\end{pmatrix}\begin{pmatrix}
\sigma_x^2 & 0 & 0 \\
0 & \sigma_v^2 & 0 \\
0 & 0 & \sigma_a^2
\end{pmatrix}\begin{pmatrix}
1 & 0 & 0 \\
0 & dt & 1 \\
0 & dt & 1
\end{pmatrix}
\]
**Example:** Tracking of a Rocket with a Kalman Filter

Propagate Error Covariance Matrix:  \[ P_{i+1} = M P_i M^T \]

\[
P_0 = \begin{pmatrix}
\sigma_x^2 & 0 & 0 \\
0 & \sigma_v^2 & 0 \\
0 & 0 & \sigma_a^2 \\
\end{pmatrix}
\]

At the next time step:

\[
P_1 = \begin{pmatrix}
\sigma_x^2 + \sigma_v^2 dt^2 & \sigma_v^2 dt & 0 \\
\sigma_v^2 dt & \sigma_v^2 + \sigma_a^2 dt^2 & \sigma_a^2 dt \\
0 & \sigma_a^2 dt & \sigma_a^2 \\
\end{pmatrix}
\]
The Rocket

Position

Velocity

Acceleration
Kalman Filter has specified the external Forcing

Forcing is specified based on the Dynamics provide by the physical Model
Next, consider the more complicated situation:

- Much more complicated Differential Equations
- Global Reconstruction
- Many observations
- Different kinds of instruments measuring different quantities
- Observations are in different places

⇒ This is the Situation in the Ionosphere
Another Data Assimilation Technique

The Kalman Filter

This is computationally the most expensive step in the Kalman filter

\[
\begin{align*}
  x^f &= Mx + \eta \\
  P^f &= MPM^T + Q \\
  y^o &= Hx + \epsilon \\
  K &= P^fH^T(HP^fH^T + R)^{-1} \\
  x^a &= x^f + K(y^o - Hx^f) \\
  P^a &= (I - KH)P^f
\end{align*}
\]
Ways to get around the Problem

→ Approximate Kalman Filters

• Band-Limited Kalman Filter
• Reduced State Kalman Filter
• Gauss-Markov Kalman Filter
• Ensemble Kalman Filter

Do not evolve Error Covariance Matrix with Model

INSTEAD

Obtain Error Covariance Matrix from an ENSEMBLE of Model runs
Gauss-Markov Kalman Filter Model (GAIM-GM)

Specification & Forecast of the Global Ionosphere

• Ionospheric Forecast Model provides background densities
• Kalman filter solves for derivations from the background
• Uses simple statistical model instead of full physics
• Error covariances are calculated from 1104 IFM model runs

• Assimilates 5 data types:
  - Slant TEC from ground-based GPS receivers
  - Bottomside $N_e$ Profiles from Ionosondes
  - UV radiances (1356Å and 911Å)
  - DMSP IES in situ $N_e$
  - Slant TEC from COSMIC
GAIM-GM Model Run for November 20, 2003 Storm
Illustration of Locations of GPS/TEC Data. Slant TEC Values have been mapped to the Vertical Direction.

GAIM Specification of TEC Distribution
Illustration of Locations of GPS/TEC Data. Slant TEC Values have been mapped to the Vertical Direction.

GAIM Specification of TEC Distribution

2003/324 20:00
Full Physics Kalman Filter Model

- **Ensemble Kalman Filter**
  - 30 Global Simulations are Launched at Each Assimilation Time Step

- **Physics-based Ionosphere-Plasmasphere Model**

- **Model Physics is embedded in Kalman filter**

- **Same 5 Data Sources as Gauss-Markov Model**

- **Provides both specifications for the ionospheric plasma densities and drivers.**
Determination of Ionospheric Drivers Using The Full Physics-Based GAIM Model

- Ionospheric Sensitivities to Drivers are embedded in the Covariances and are automatically and at each Time Step calculated.

- Drivers include:
  - Electric Fields
  - Neutral Wind
  - Composition
  - ...
Example of Full Physics-Based Kalman Filter Model

• Several Days in March/April of 2004

• Geomagnetically Quiet Period

• Data Assimilated
  o Slant TEC from 162 GPS Ground Receivers

• Use Ionosonde Data for Validation
Comparison with Ionosonde Data

Ionosonde Data were NOT assimilated!
Data Issues

• Are There Enough Data?
• What is the Data Quality?
• Are Error Estimates Available?
• Are Data Available in Real Time?
• Are Different Data Types Required?
How Does Missing or Incomplete Physics Affect the Data Assimilation Results?

- Simulate the Ionosphere with the IPM
- Modify the Simulated Ionosphere to Account for Missing Physics
- Generate Synthetic Data from Real Locations
- Reconstruct the Ionosphere with the Gauss-Markov Data Assimilation Model
- Compare Reconstructed Ionosphere with Original Ionosphere
Gauss-Markov Reconstruction
With Synthetic Data

Results Look Reasonable, but are Wrong

Ionosphere That Produced Synthetic Data

Four Bubbles
Summary

NCEP Operational Forecast Skill
36 and 72 Hour Forecasts @ 500 MB over North America
[100 ^ (1-S/70) Method]

- 36 Hour Forecast
- 72 Hour Forecast